

R&D at IPAG

Cristal team



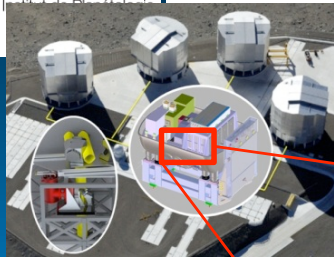
It is not an overview.

The talk is focused on

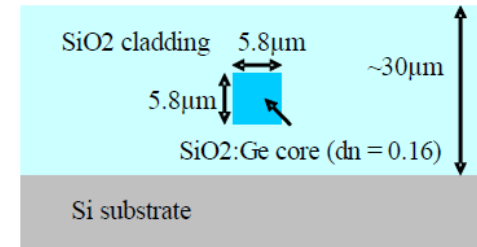
- integrated optics (combiner, spectrograph)
- detector developments



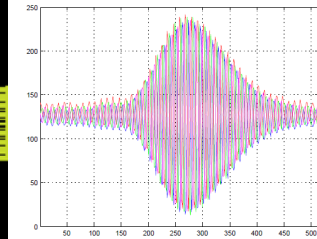
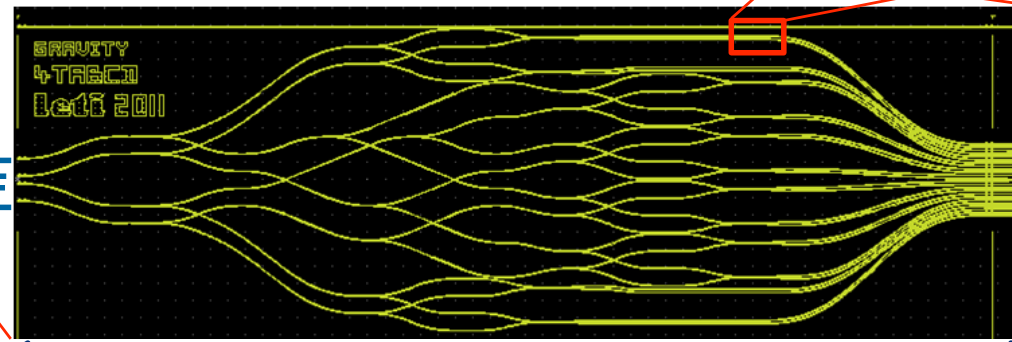
Integrated optics combiners at the heart of GRAVITY



Single-mode waveguides
Silica-on-silicon etching technique
(Collaboration CEA-LETI/IPAG)



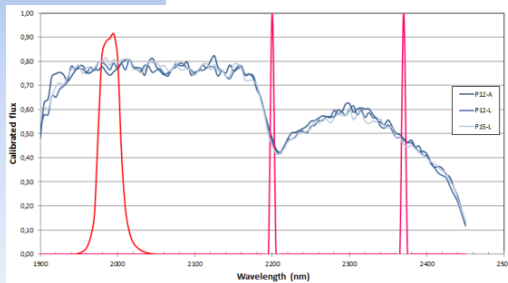
Telescopes
1 2 3 4



(ii)

[Jocou et al. 2012]

58 mm



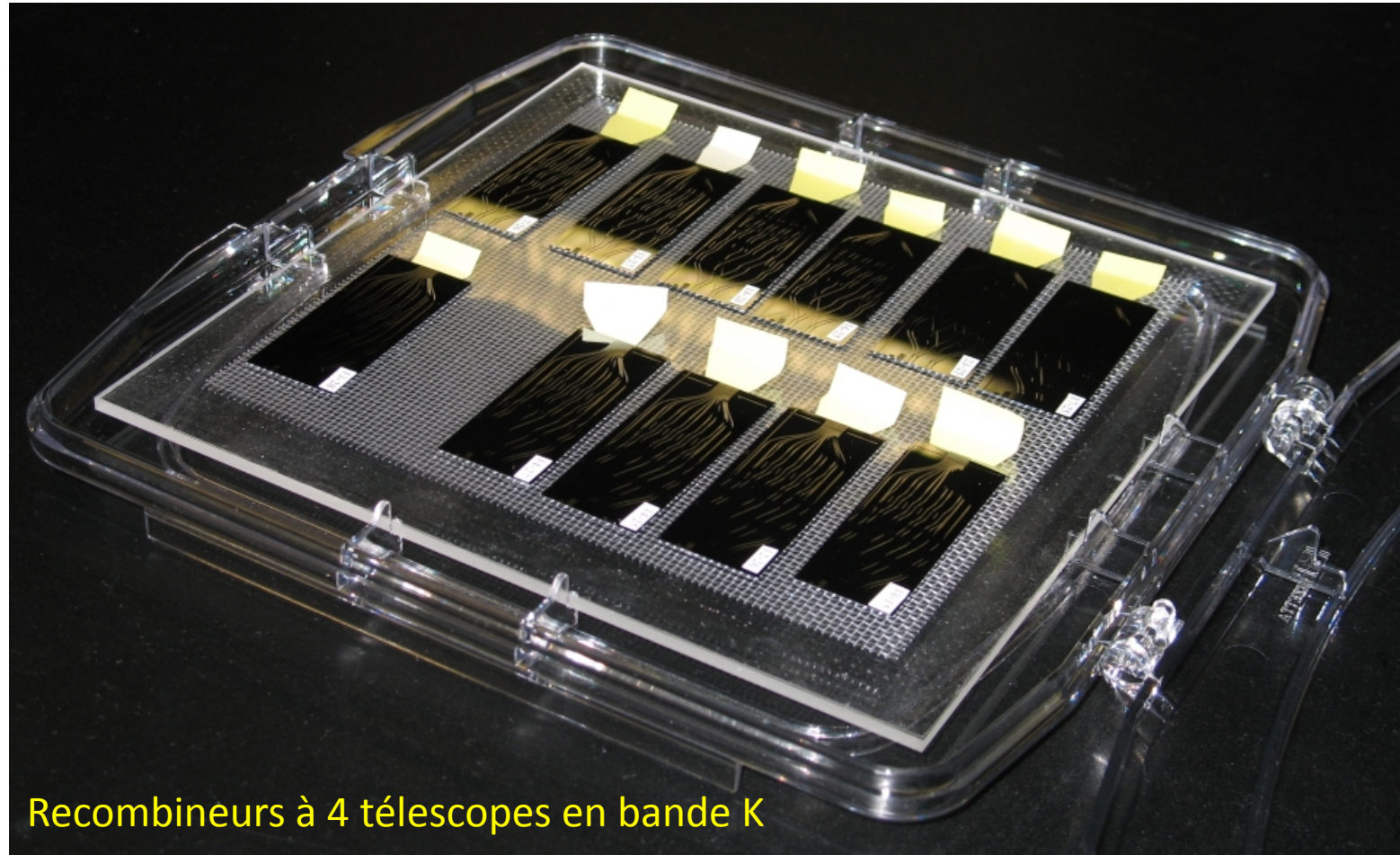
(i)

The beam combiners

- Work in the K band (i)
- Ensure the fringe coding on 6 baselines (ii)
- Operate at 200 K in vacuum
- Allows the metrology laser to be injected

Technological challenge: reach a throughput of 55% over the [2 μm ; 2.45 μm] spectral band

Final GRAVITY combiners

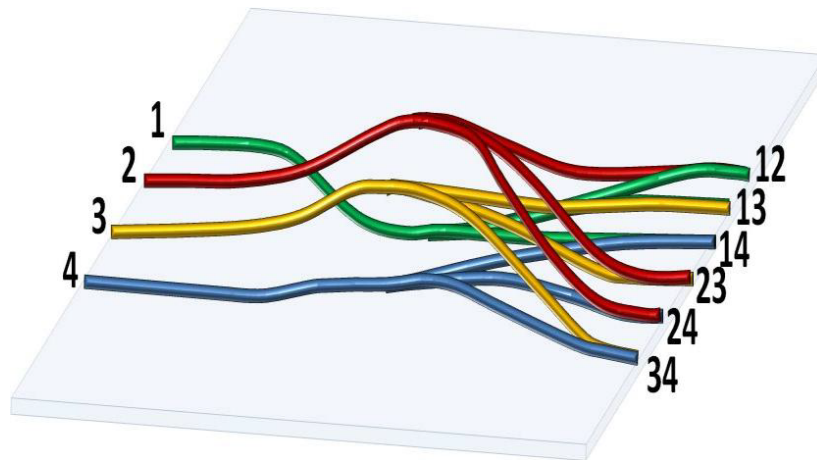
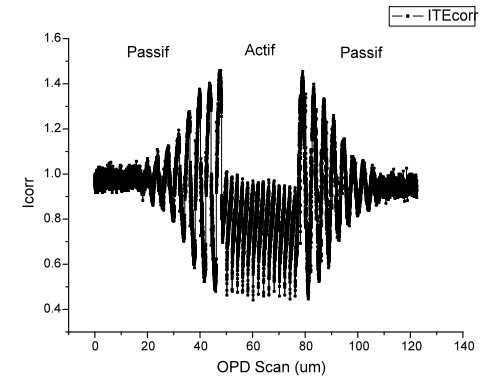


Recombineurs à 4 télescopes en bande K

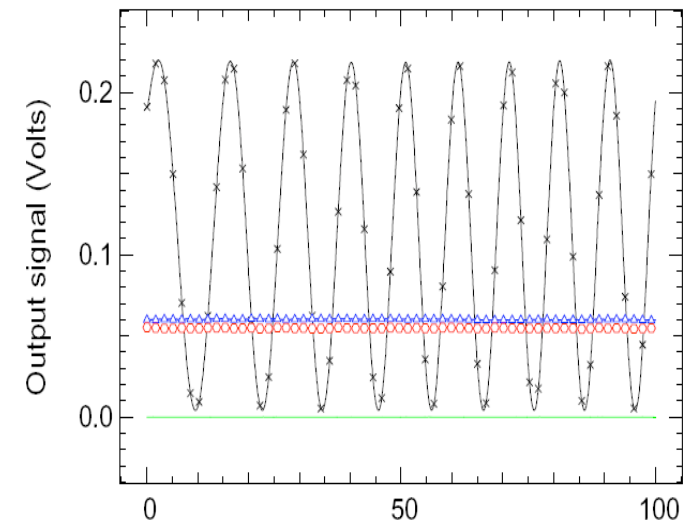
Optique Intégrée Bande L

Préparer toutes les fonctions clé pour un futur instrument en bande L

- Guides Droits, Coupleurs, Y en Bande L
- Electro-optique (modulation interne de la phase)
 - Déplacement de 4 franges @ 3.39 μm avec 100 V
- Transmission ($T(1\text{cm})=93\%$)
- Différents matériaux utilisés :
 - Niobate (@ 3-5 μm)
 - Chalcos (@ 5-15 μm)
- Design 2D, 3D (pour raccourcir les composants)



Collaboration avec :
 - PNNL (Washington)
 - Université Heriot-Watt



[Labadie et al. 2011]

Optique Intégrée Bande L

**FEMTO-ST
PHOTLINE
Besançon**

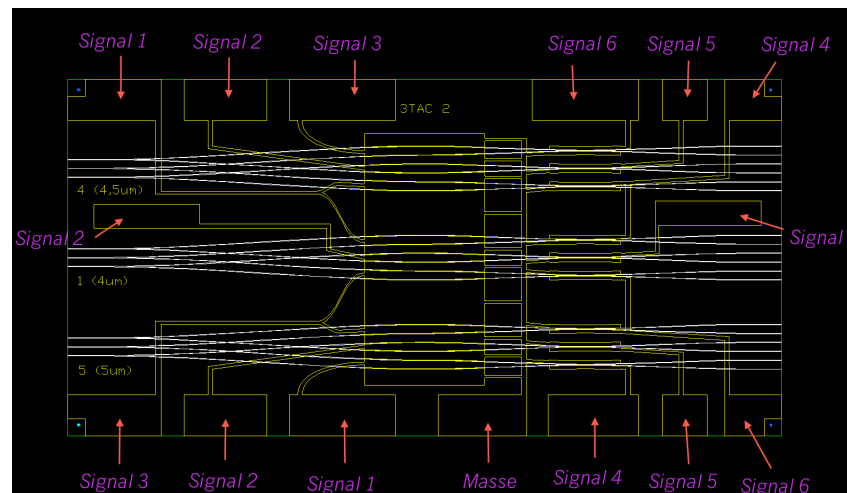
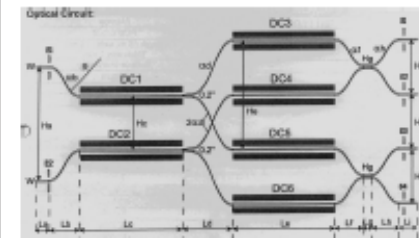
Waveguide building method :

Titanium diffusion

-20h diffusion @ 1050°C

-Both polarizations are guided

-Single mode behavior



Recombineur à 3 télescopes en bande L

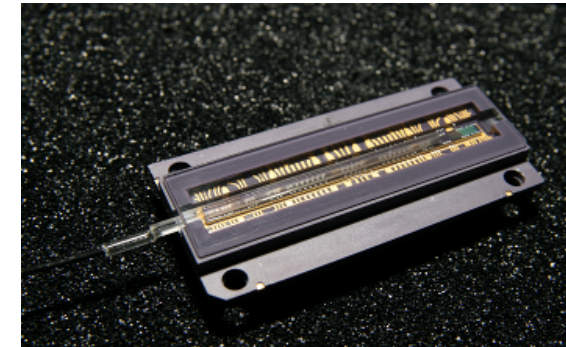
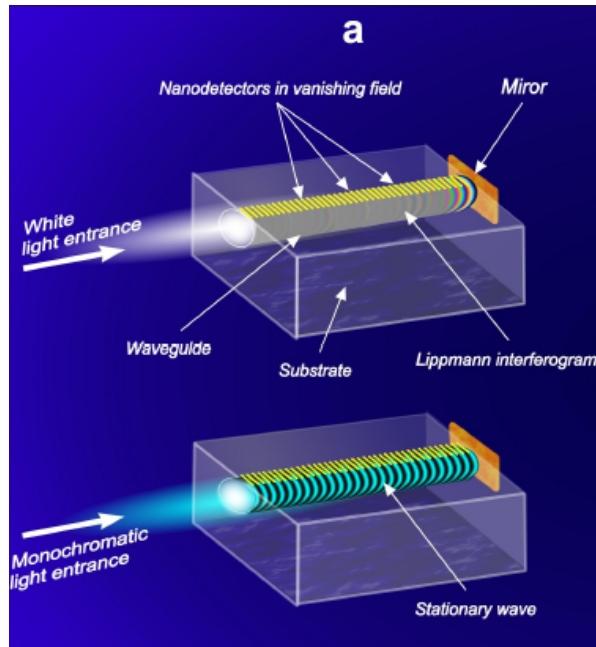
[Heidmann et al. 2012]

Spectrométrie intégrée

Idee :

Interférence dans un guide + détecteur in situ = spectromètre SWIFTS

[Le Coarer et al. 2007]



λ : [400 nm ; 1000 nm]
 R ~ 50000

- Collaboration large : laboratoires académiques et partenaires industriels (FUI)
- Valorisation: start-up Resolution Spectra Systems
- IPAG: concepts, études système, calibration, traitement des données, tests



Detectors for astronomy

- ❑ **Connecting and stimulating various actors in « detector » world:**
= Labex « FOCUS » started in 2012 (PI P. Kern)

- ❑ **Long wavelengths:** towards high survey efficiency (numerous and sensitive broadband pixels) based on KIDS.

- ❑ **Optical and NIR:** fast and sensitive sensors
 - ✓ **RAPID** *Sofradir – BiospaceLab – LAM – IPAG – CEA/LETI – ONERA*
Based on avalanche photodiodes

 - ✓ **OCAM** *FP6/Opticon, E2V, OHP, LAM, IPAG, ESO*
Based on EMCCD



Based on avalanche photodiodes
Funded by a FUI (2009-2012)

Today:

4 chips available

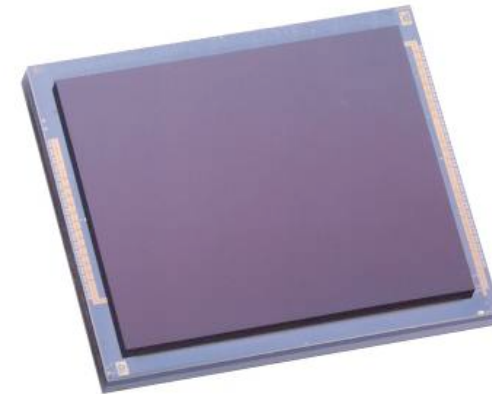
2 are integrated in its cryostat (at IPAG)

Both cameras are equipped with a cold
glued H-filter and operate at -200°C
(without nitrogen)

The first camera will be soon delivered to
ONERA for electronics characterization

First images by 2 months

Future: go to the K band (new filter)



320x256 pixels (30 μm)
 λ : 0.4-3 μm
1500 images/s
8 outputs
Noise: 3-4 e⁻ (gain = 10) measured
by CEA/LETI

OCAM

Based on EMCCD (E2V detector)

Funded by OPTICON

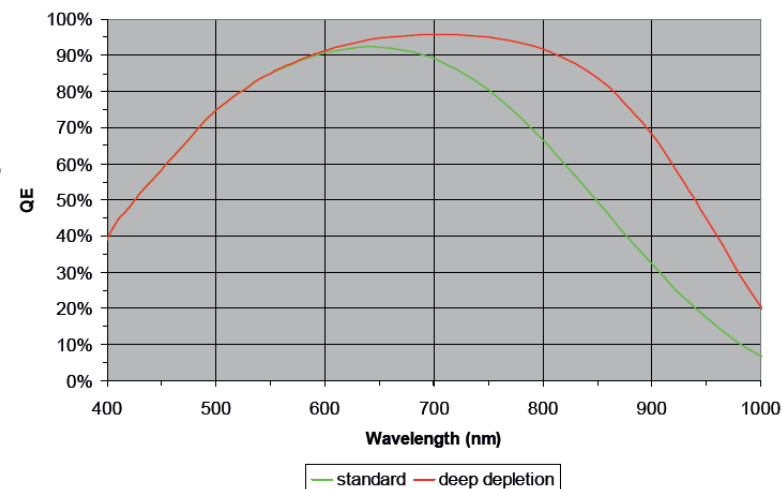
Designed to match the SPHERE requirements



240x240 pixels (24 μm)

Noise: 0.5e- @ 1500Hz

CCD220 QE at -45°C Basic ER1



NOISE HISTOGRAM FOR GAIN X1000

