

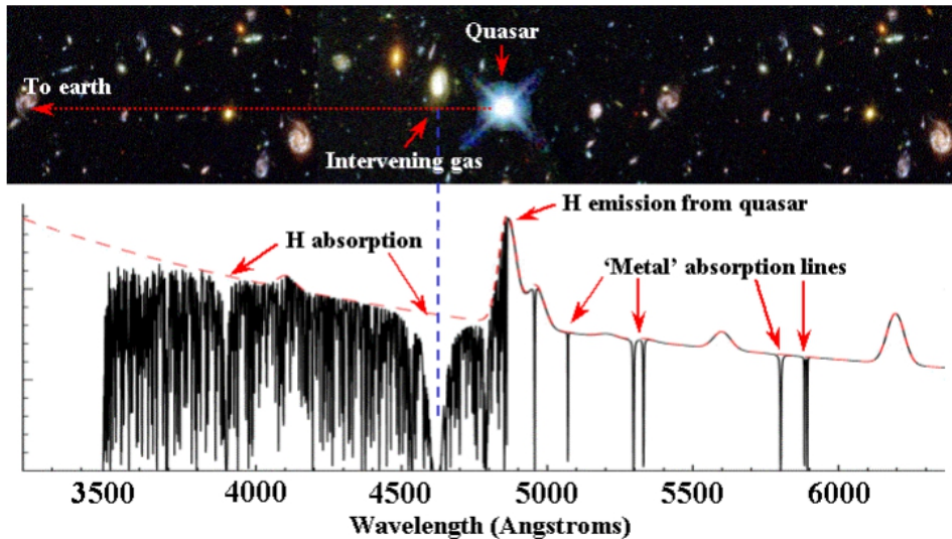
The Chemical connection between Damped Lyman-alpha systems and dwarf galaxies

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DLAs vs Local Dwarf Galaxies

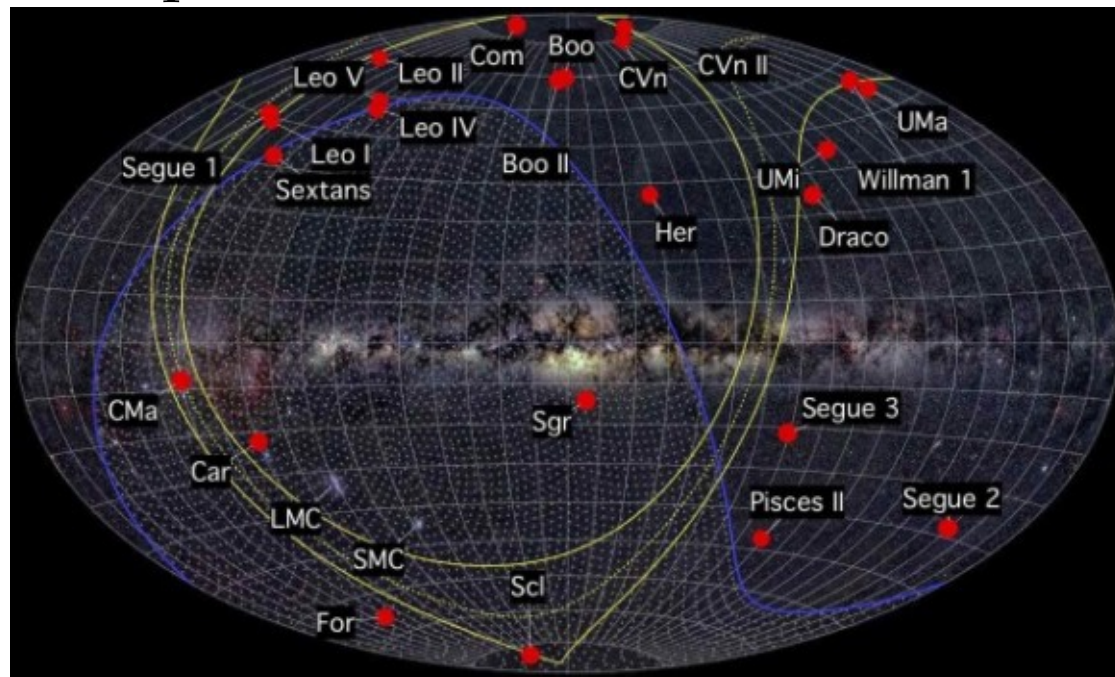


- ★ DLA systems are reservoirs of neutral gas ($N(\text{H I}) \geq 2 \times 10^{20} \text{ cm}^{-2}$), visible due to absorption.
- ★ Dust depletion!
- ★ Volatile elements such as S, Zn, O important.

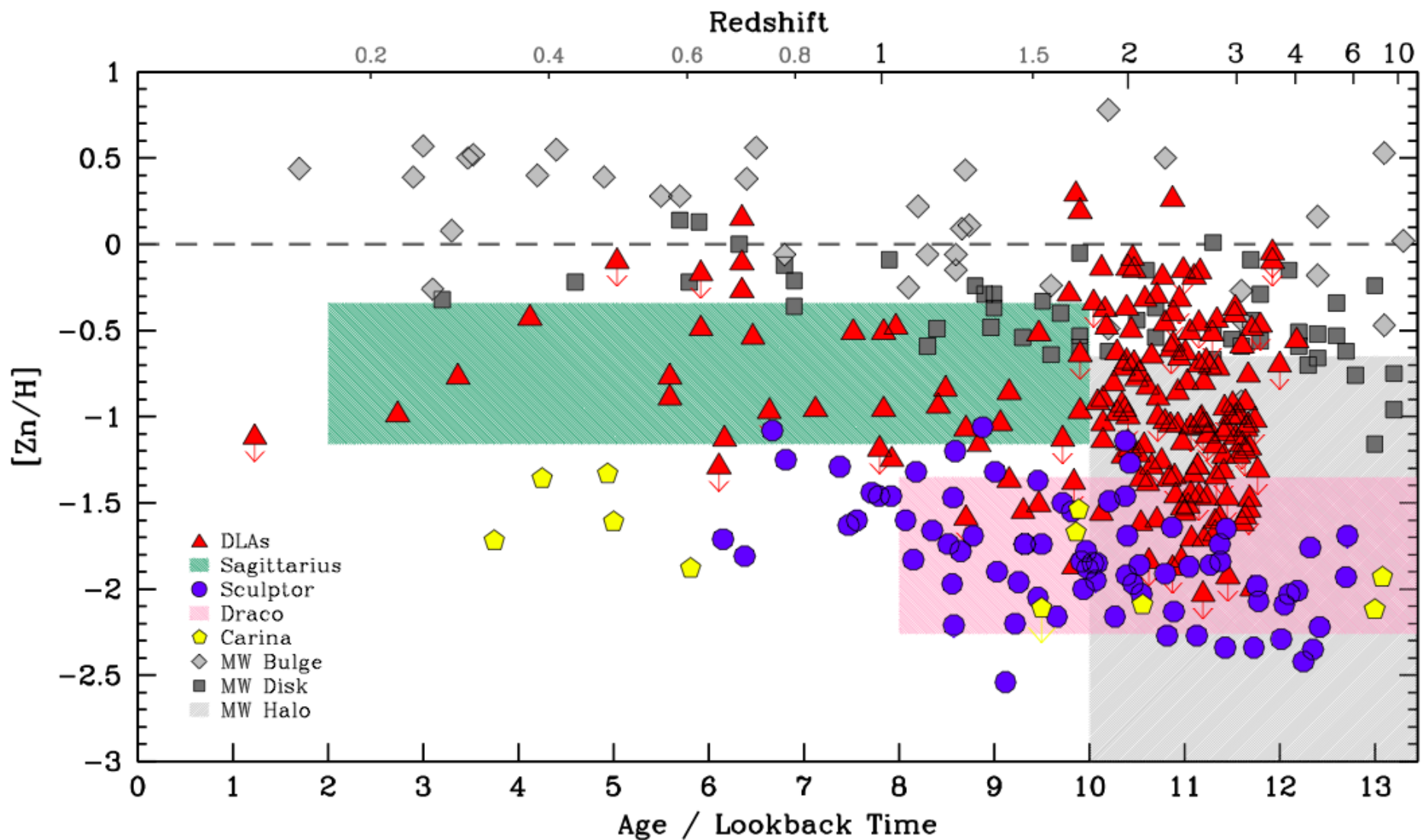
★ In the Local Group we are able to observe individual stars and get detailed chemical abundances.

★ 50+ known dwarf galaxies around the Milky Way

★ Stellar physics!



Metallicity Comparison

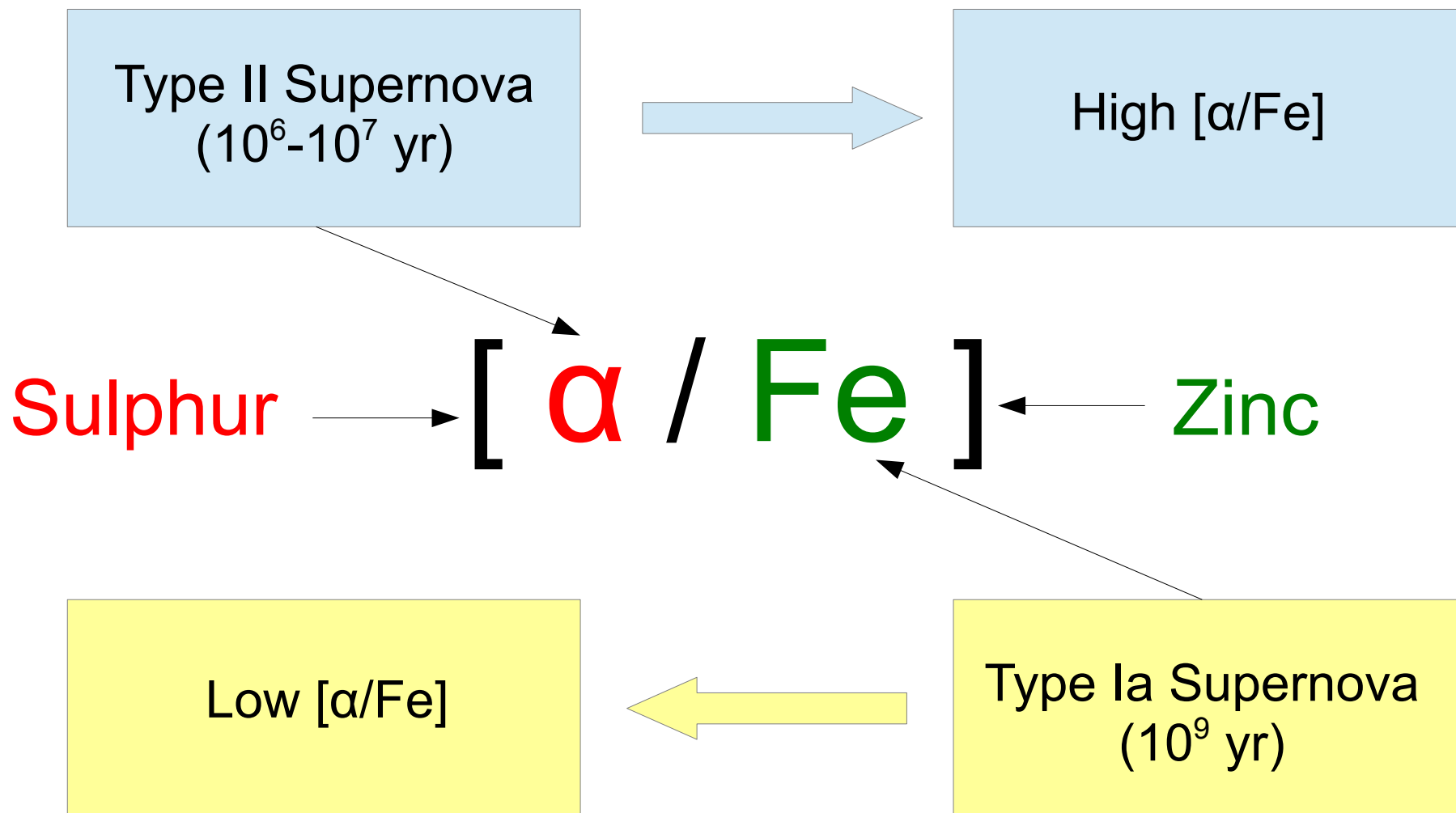


DLAs: Quieret+16; Milky Way: Cayrel+04, Reddy+06, Nissen+07, Bensby+2013;

Sculptor: Skúladóttir+17, deBoer+12; Carina: Shetrone+03, Venn+12, Lemasle+12; Sagittarius: Sbordone+07.

Age models: Salvadori+10, Aparicio+11, Bellazzini+06, Siegel+07, deBoer+15

Abundance ratios



Abundance ratios

[α / Fe]



[S / Zn]

- ★ Never been compared with stellar dwarf galaxy abundances
- ★ **Now for the first time 80+ stars in Sculptor!**

(Skúladóttir+ 2015b; 2017a)

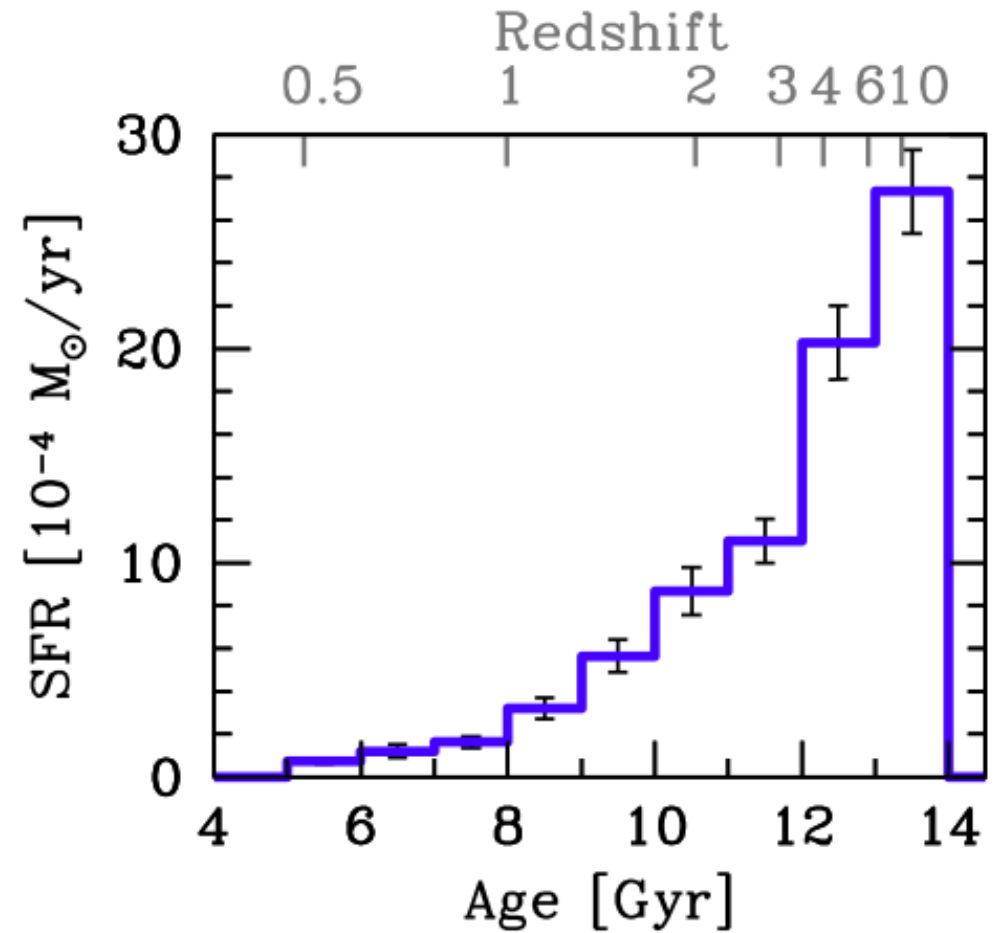
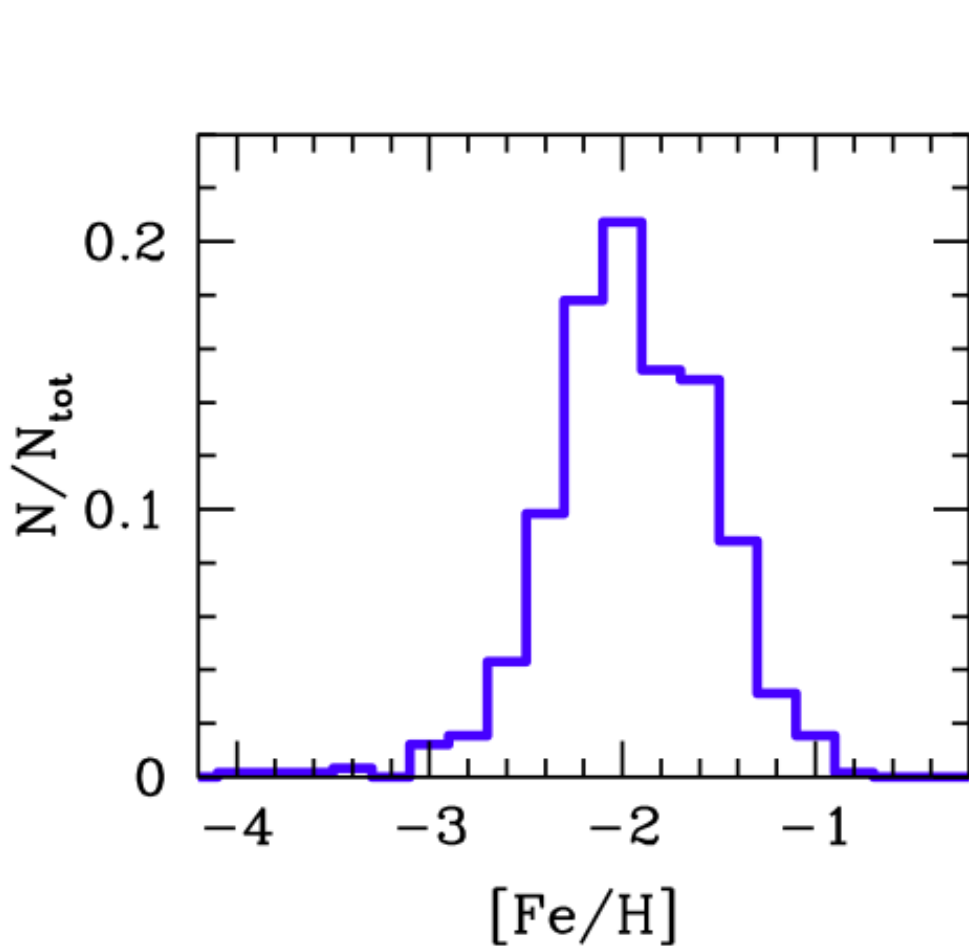
The Sculptor dwarf spheroidal

- ★ Discovered in 1938
- ★ $M_{\text{tot}} = 3.4 \times 10^8 M_{\odot}$
- ★ $M_{\text{stars}} = 8 \times 10^6 M_{\odot}$
- ★ Distance of 86 ± 5 kpc
- ★ At high Galactic latitude



The stellar population in Sculptor

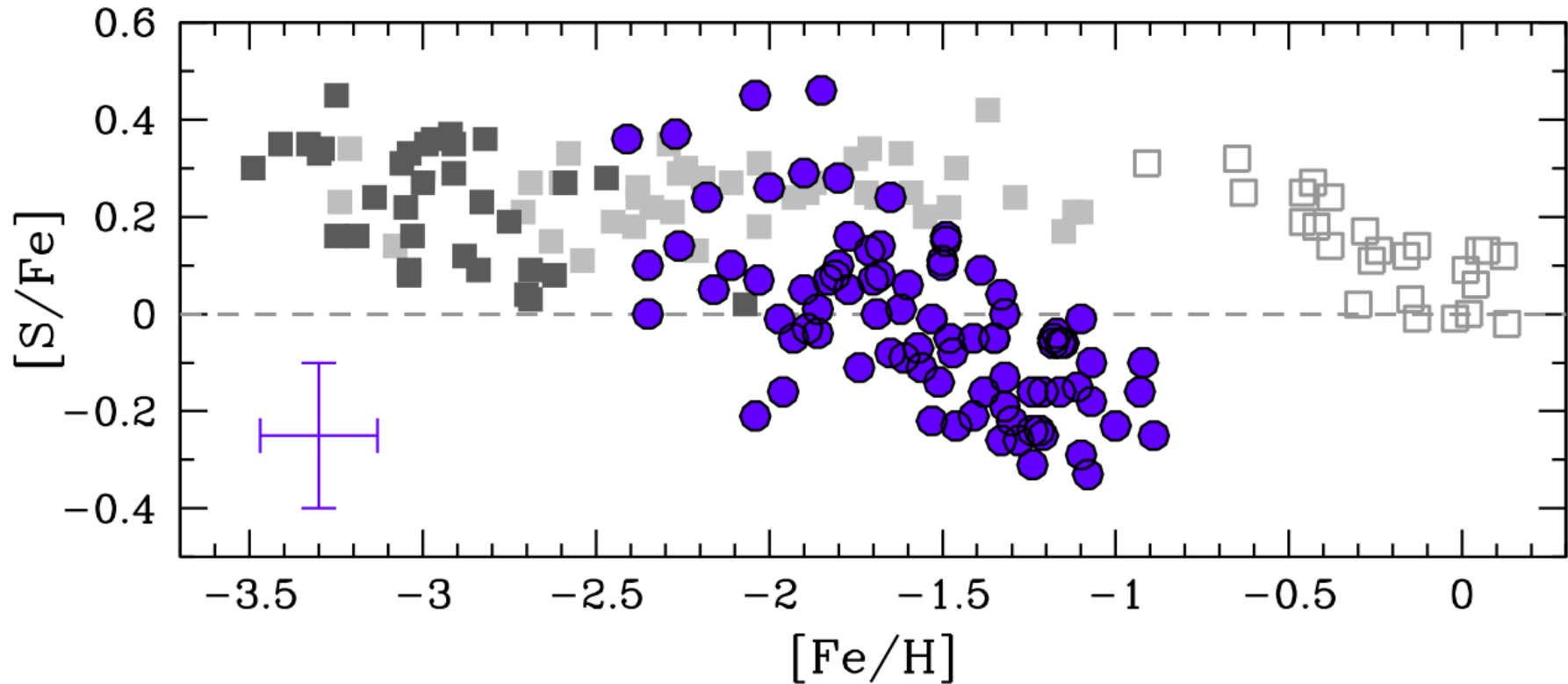
★ Dominated by old stars (>10 Gyr old)



Metallicity Distribution Function
(Starkenberg et al. 2010)

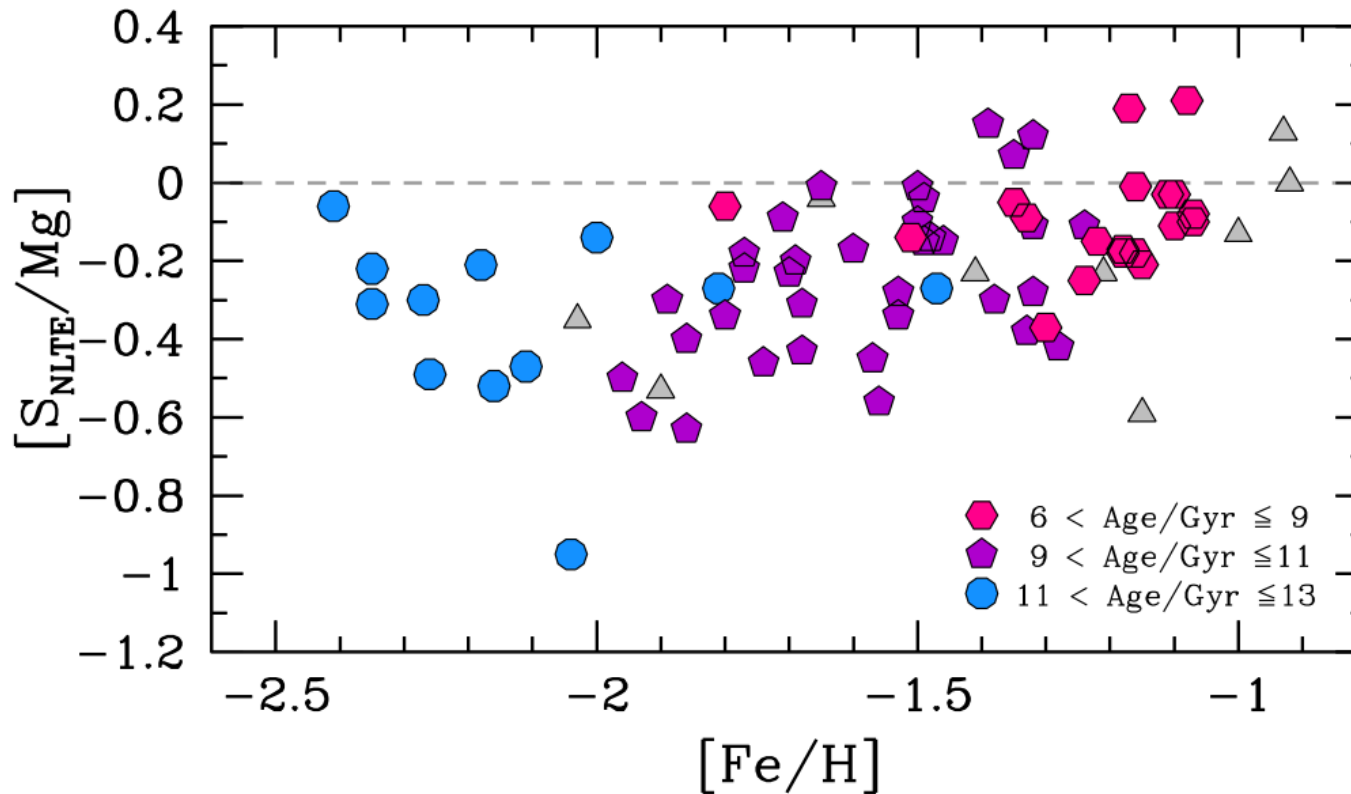
Star Formation History
(de Boer et al. 2012)

Closer Look - Sulphur



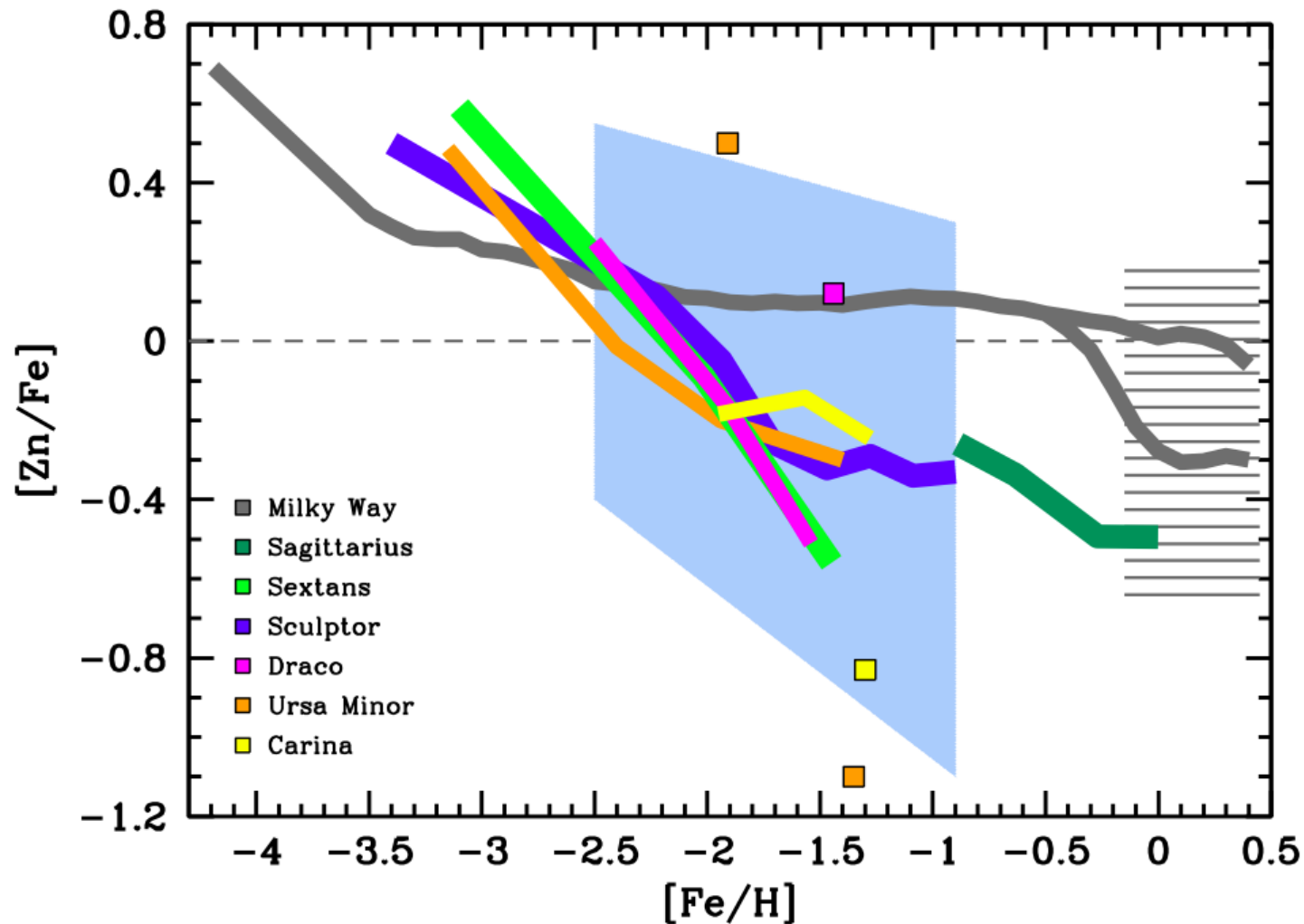
- ★ The sulphur abundances in Sculptor show the same behaviour as other α -elements, consistent with a plateau at lower metallicities; with increasing $[Fe/H]$, the ratio $[S/Fe]$ declines, reaching negative values at the highest $[Fe/H]$.

[S/ α] in Sculptor



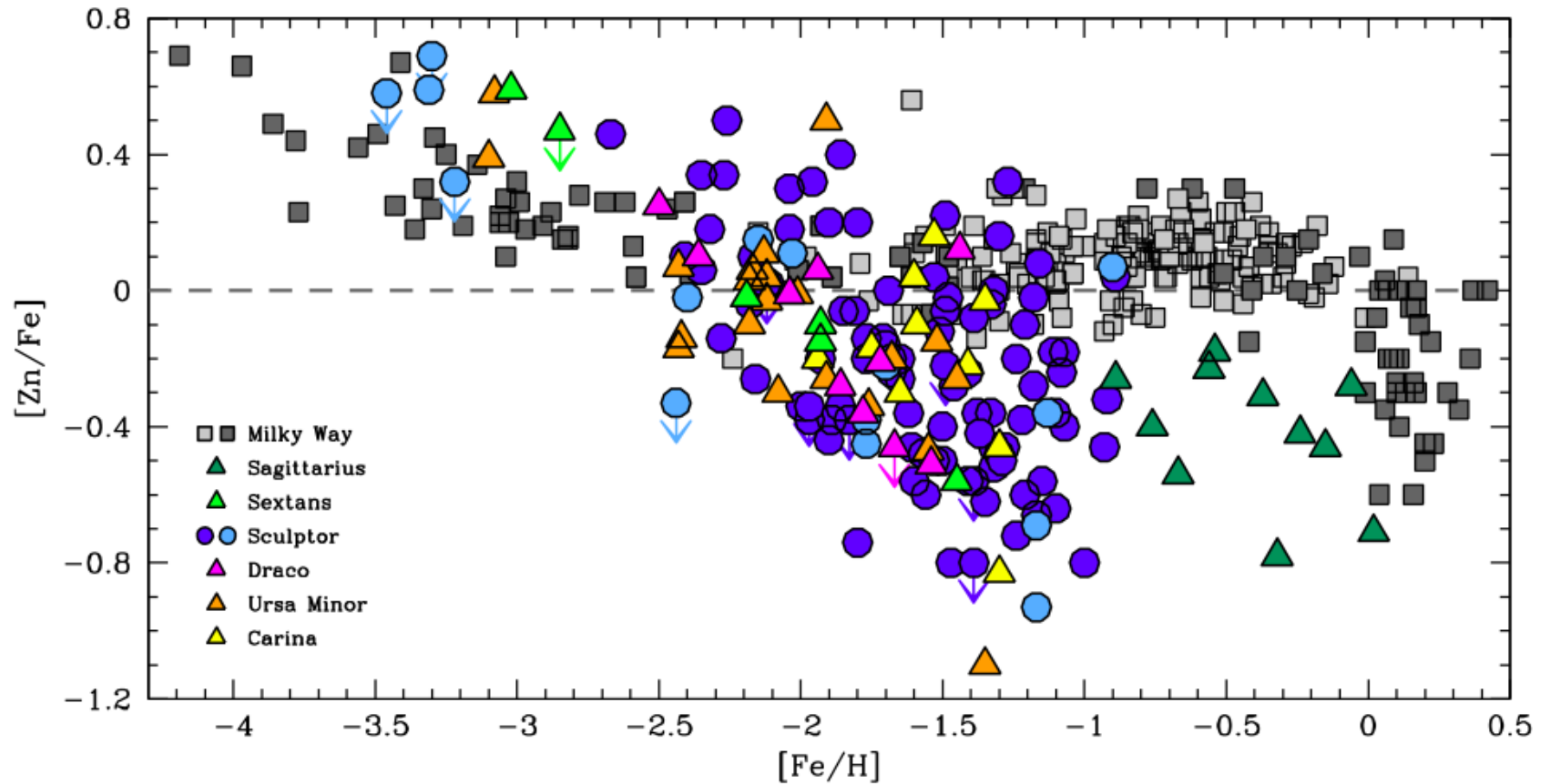
- ★ The ratio of sulphur to magnesium increases with $[\text{Fe}/\text{H}]$, consistent with **some production** of S in **Supernovae Type Ia**

Closer Look - Zinc



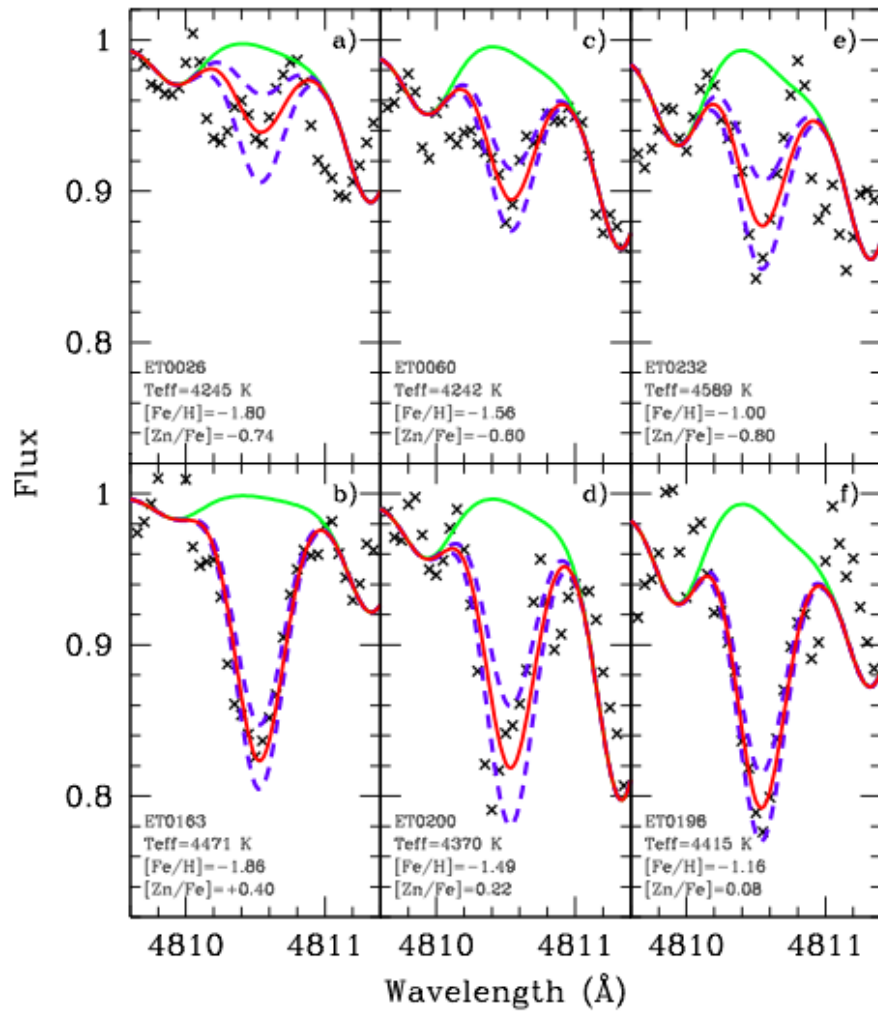
★ **Zn behaves like an α -element! But with scatter!**

Closer Look - Zinc



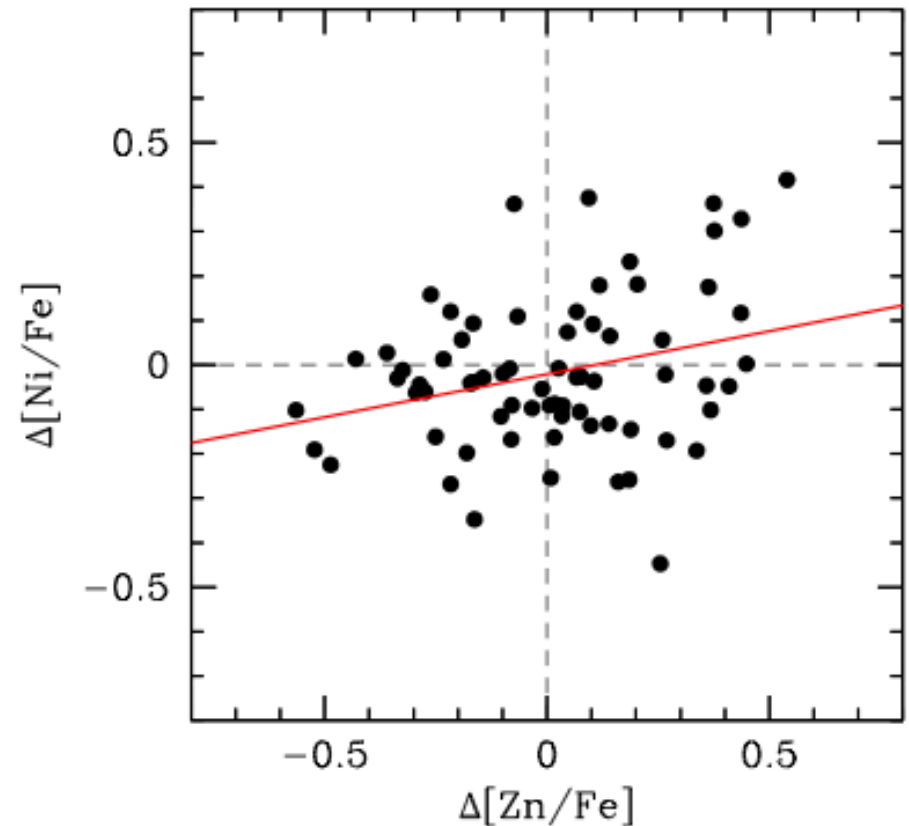
References (Number of stars): *Sagittarius*: Sbordone et al. 2007 (11). *Sextans*: Shetrone et al. 2001 (5); Honda et al. 2011 (1). *Sculptor*: Shetrone et al. 2003 (1); Geisler et al. 2005 (1); Kirby & Cohen 2012 (1); Jablonka et al. 2015 (4 upper limits); Skúladóttir et al. 2015b (1); Simon et al. 2015 (1); Hill et al. in prep (6). *Draco*: Shetrone et al. 2001 (5); Cohen & Huang 2009 (5). *Ursa Minor*: Shetrone et al. 2001 (6); Sadakane et al. 2004 (3); Cohen & Huang 2010 (10); Ural et al. 2015 (2). *Carina*: Shetrone et al. 2003 (5); Venn et al. 2012 (5). *Milky Way*: Reddy et al. 2003, 2006; Cayrel et al. 2004; Nissen & Schuster 2011; Ishigaki et al. 2013; Bensby et al. 2014 (only including stars with errors $\delta_{[Zn/Fe]} \leq 0.2$); Barbuy et al. 2015.

Scatter

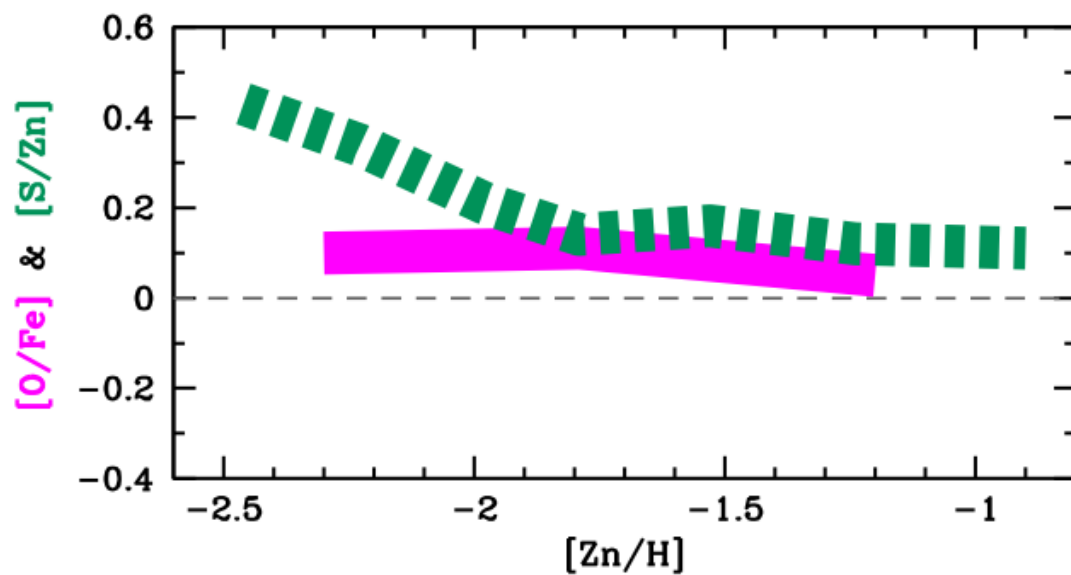
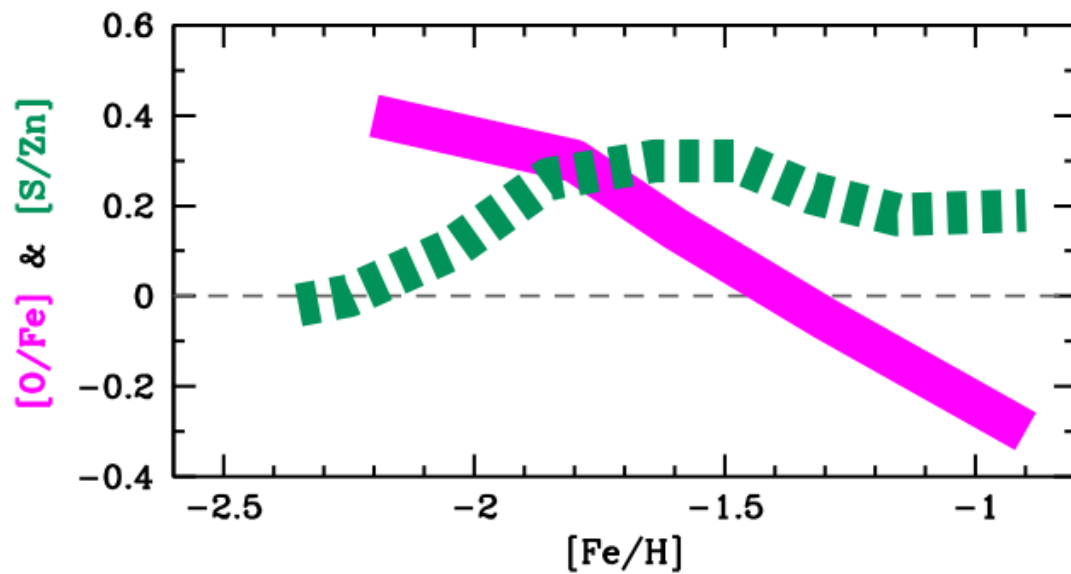


★ Scatter in $[\text{Zn}/\text{Fe}]$ correlated with Ni

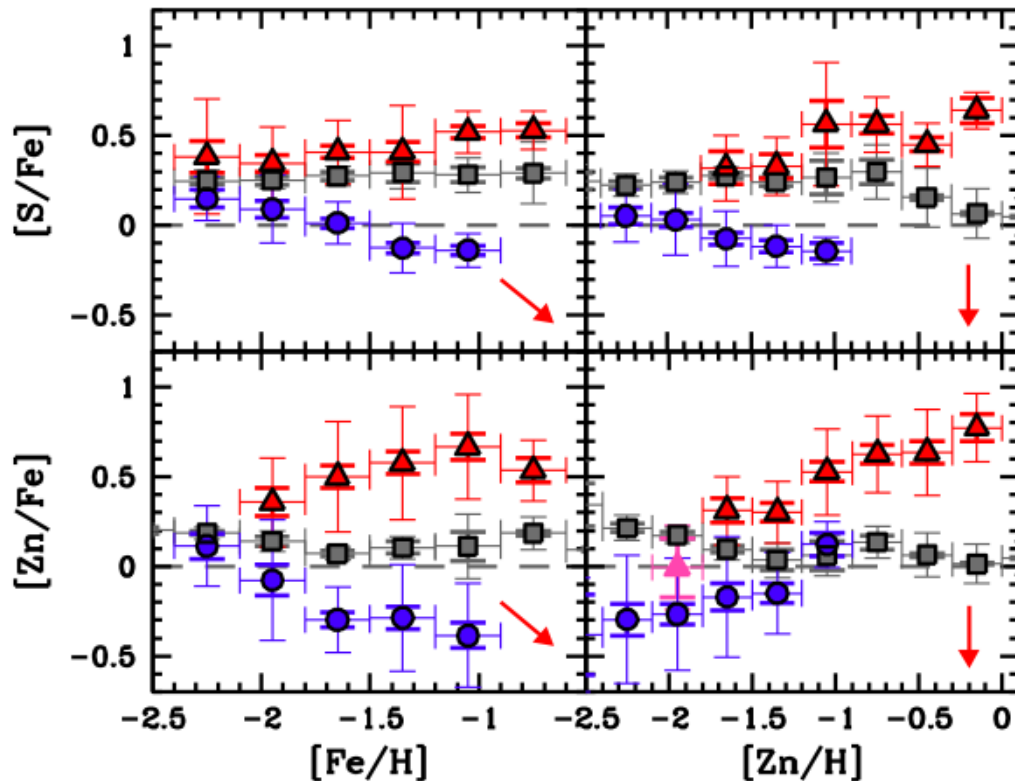
★ (but not: O, Na, Mg, Si, Ca, Ti, Sc, Cr, Mn, Fe II, Co, Ni, Ba, La and Nd)



$[S/Zn] \neq [\alpha/Fe]$

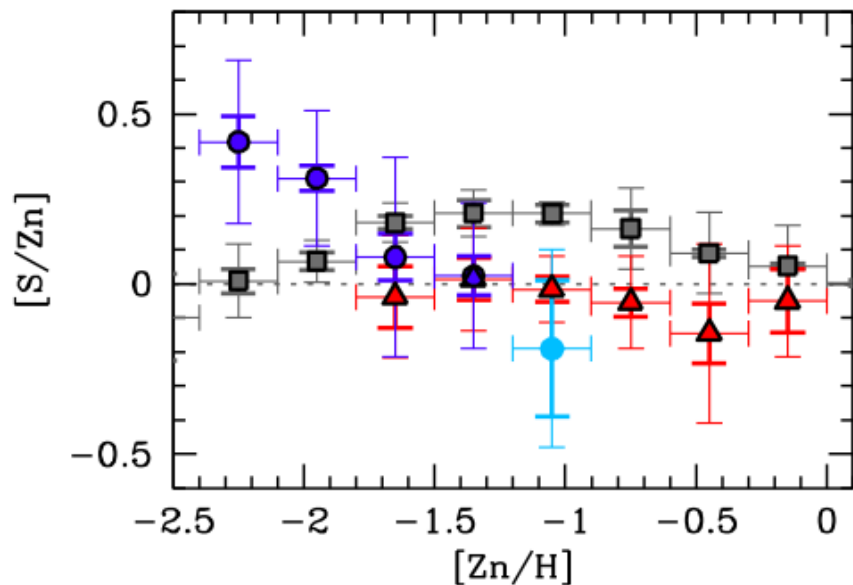
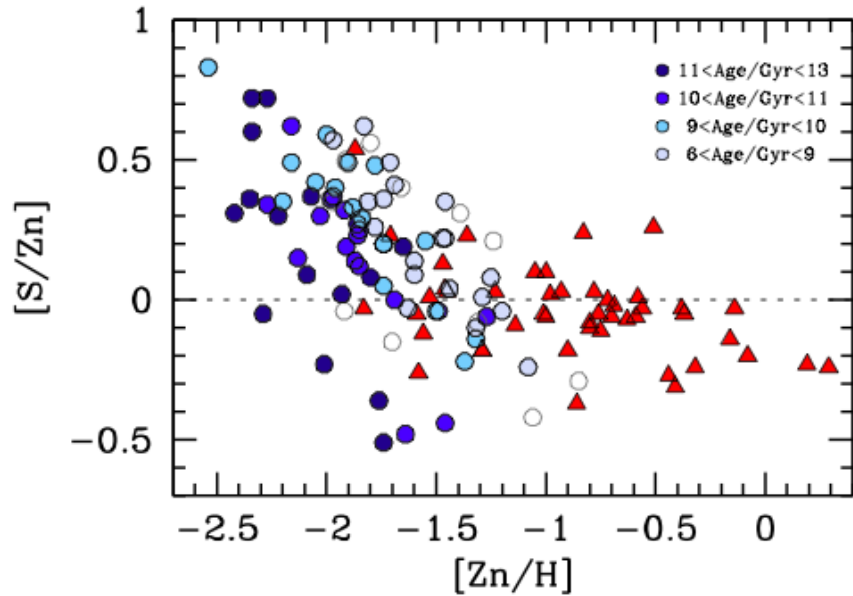


DLAs vs Sculptor



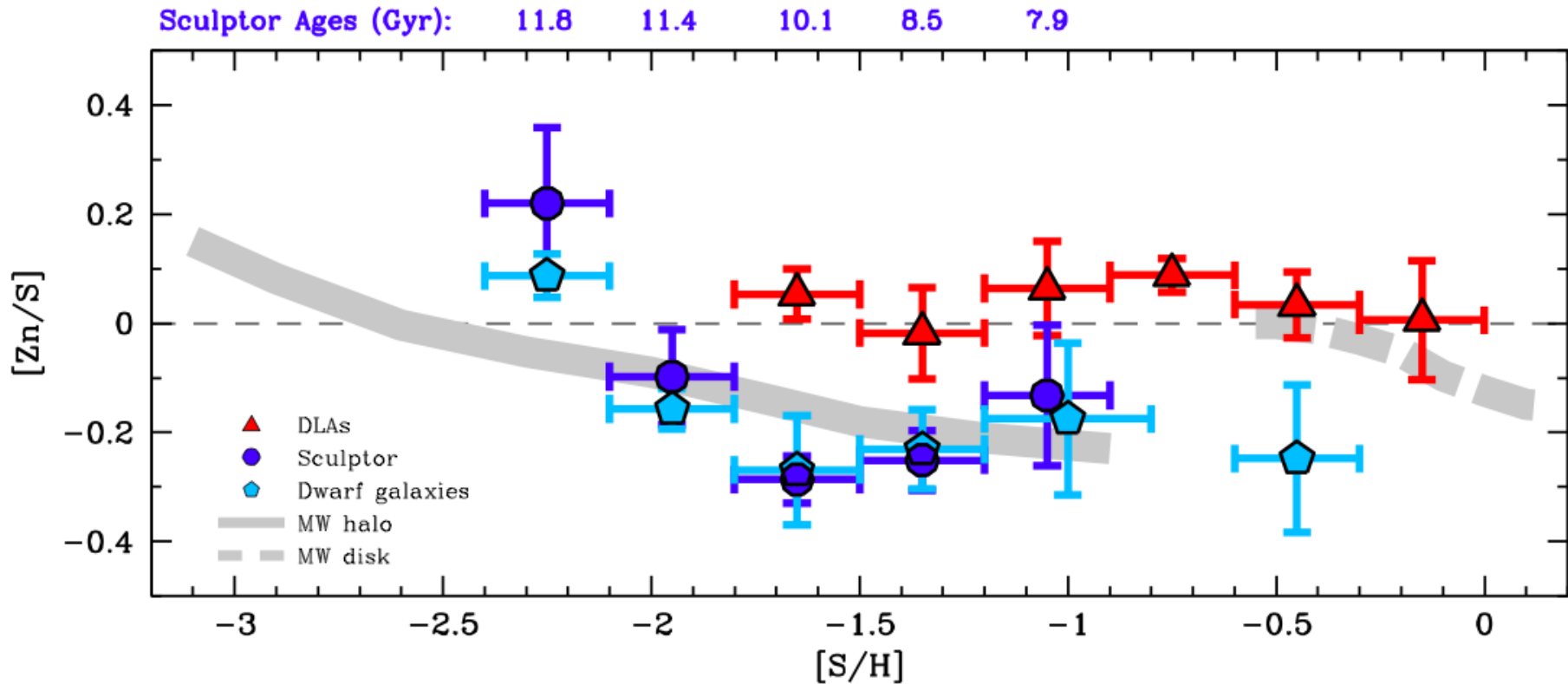
- ★ Sculptor stars cover lower $[Zn/H]$ compared to DLAs
- ★ DLAs, Milky Way and Sculptor all different!
- ★ Dust depletion of Fe complicates matter.
- ★ In principle, DLAs could be like the Milky Way, or Sculptor, or neither depending on the level of dust depletion.

DLAs vs Sculptor



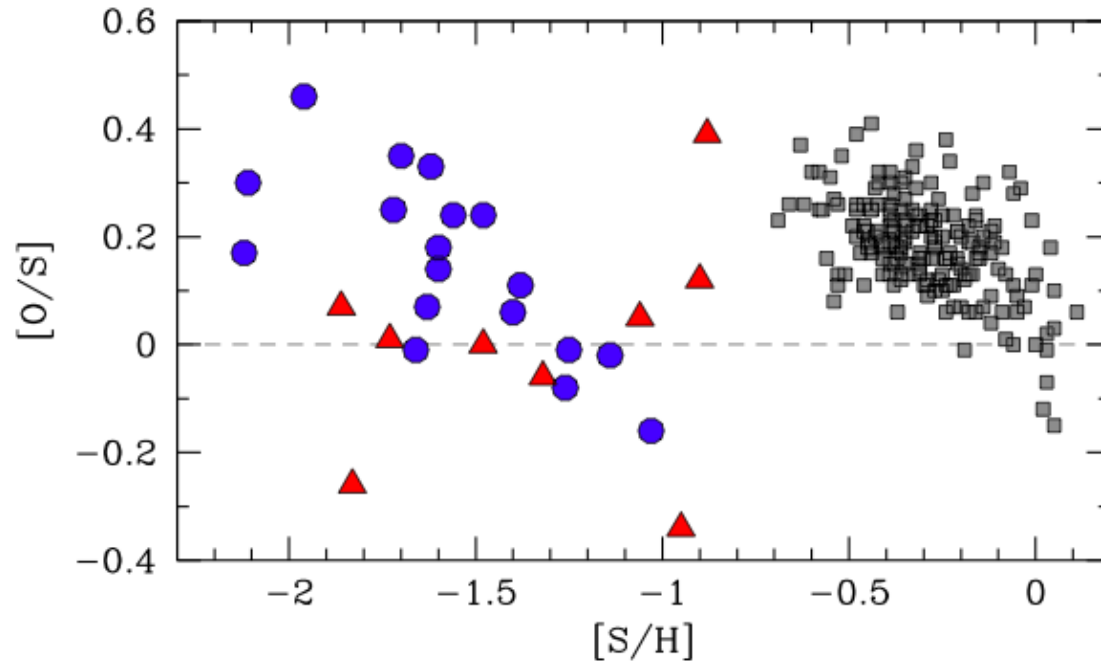
- ★ $[S/Zn]$ consistent in regions where they overlap!
- ★ However: $[Zn/H]$ is not a good tracer of the chemical evolution in Sculptor.
- ★ Results difficult to interpret within the star formation history of Sculptor

DLAs vs Sculptor



- ★ Clear differences in abundance pattern!
- ★ DLAs comparable to the early dwarf galaxy evolution, and the disk at higher metallicities.
- ★ The plateau in DLAs a mystery!

Alternative $[\alpha/\text{Fe}]$: $[\text{O}/\text{S}]$



- ★ Instead of $[\alpha/\text{Fe}]$ with $[\text{Fe}/\text{H}]$ we propose the use of $[\text{O}/\text{S}]$ with $[\text{S}/\text{H}]$ as a measure of the relative contribution of SN Type II and Type Ia.
- ★ Observationally challenging in DLAs.

Conclusions

★ Zn is not Fe

- ★ Contrary to Fe, Zn is not significantly created in Supernova Type Ia.
- ★ $[S/Zn] \neq [\alpha/Fe]$
- ★ Zn is still not fully understood in the Local Group.
- ★ S can be used as a metallicity tracer, created both by SN Type II and Ia.
- ★ DLA abundance ratios of $[Zn/S]$ consistent with earliest phases of dwarf galaxy evolution.
- ★ Plateau of $[Zn/S]$ in DLAs a mystery!